

Based on Electromagnetic Radiation				
Type Spectroscopy	Usual Wavelength Range*	Usual Wavenumber Range, cm <sup>-1</sup>	Type of Quantum Transition	
Gamma-ray emission	0.005–1.4 Å	-	Nuclear	
X-Ray absorption, emission, fluorescence, and diffraction	0.1–100 Å	_	Inner electron	
Vacuum ultraviolet absorption	10–180 nm	$1\times 10^6$ to $5\times 10^4$	Bonding electrons	
Ultraviolet visible absorption, emission, and fluorescence	180–780 nm	$5\times10^4$ to $1.3\times10^4$	Bonding electrons	
Infrared absorption and Raman scattering	0.78–300 µm	$1.3\times10^4$ to $3.3\times10^1$	Rotation/vibration of molecules	
Microwave absorption	0.75–3.75 mm	13–27	Rotation of molecules	
Electron spin resonance	3 cm	0.33	Spin of electrons in a magnetic field	
Nuclear magnetic resonance	0.6–10 m	$1.7\times 10^{-2}$ to $1\times 10^{3}$	Spin of nuclei in a magnetic field	

Source: Skoog, Holler, and Nieman, Principles of Instrumental Analysis, 5th edition, Saunders College Publishing.





### **Wave Properties**

<u>Period</u> (p) – the time required for the passage of successive maxima through a fixed point in space.

<u>Frequency</u> (v) – the number of oscillations of the field that occur per second. Equal to 1/p. Determined by source and remains invariant regardless of media traversed.

<u>Velocity</u> (v) – the rate at which a wave front moves through a medium. Dependent on composition of medium and frequency.

### **Wave Properties (continued)**

<u>Wavelength</u>  $(\lambda)$  – the linear distance between successive maxima or minima of a wave. The wavelength must decrease as radiation passes from a vacuum to some other medium.

<u>Wavenumber</u> ( $\sigma$ ) – the number of waves per centimeter.





## **Principle of Superposition**

When two or more waves traverse the same space, a displacement occurs which is the sum of the displacements caused by the individual waves.

<u>Constructive vs. Destructive Interference</u> – based on phase difference between waves.

<u>Fourier transform</u> based on fact that any wave motion, regardless of complexity, can be described by a sum of simple sine or cosine terms.

<u>Diffraction</u> – process in which a parallel beam of radiation is bent as it passes a sharp barrier or through a narrow opening. A consequence of interference.



#### **Refraction of Radiation**

**<u>Refractive Index</u>**:  $n_i = c/v_i$ where *c* is the speed of light in a vacuum

Dispersion – the variation of refractive index of a substance

<u>Dispersion</u> – the variation of refractive index of a substa with frequency or wavelength.

<u>Normal Dispersion</u> – gradual increase in refractive index with increasing frequency (or decreasing wavelength).

<u>Anomalous Dispersion</u> – sharp change in refractive index is observed. Always occurs at frequencies that correspond to the natural harmonic frequency associated with some part of the molecule, atom, or ion of a substance. At these frequencies, permanent energy transfer from the radiation to the substance occurs and absorption of the radiation is observed.





## **Scattering of Radiation**

<u>Rayleigh Scattering</u> – scattering by molecules or aggregates of molecules with dimensions significantly smaller than the wavelength of radiation. Intensity related to wavelength, dimensions of scattering particles, and polarizability.

<u>Raman Scattering</u> – part of the scattered radiation suffers from quantized frequency changes as a result of vibrational energy transitions occurring in a molecule as a consequence of the polarization process.

# **Absorption of Radiation**

Selective removal of certain frequencies by transfer of energy to atoms or molecules.

Particles promoted from lower-energy (ground) states to higherenergy (excited) states.

Energy of exciting photon must exactly match the energy difference between the ground state and one of the excited states of the absorbing species.







Table 8.3 The Transformations of Light Energy Interacting with Atoms			
Energy in	Energy out	Spectrometry	
Heat	Light	Emission (incandescence)	
Light	Heat	Absorption	
Light	Light	Luminescence (phosphores- cence, fluorescence)	
Light	Moving electrons	Photoelectron spectroscopies	
Moving electrons	Moving electrons	Auger	
Bonding energy	Light	Chemiluminescence	







# **Radiation Sources**

Sufficient power Suitable stability

**Types:** 

- ≻ Continuous sources *e.g.*, lamps used for absorption
- ≻Line sources *e.g.*, vapor lamps used in AA
- **>**Lasers (light <u>a</u>mplification by <u>s</u>timulated <u>e</u>mission of <u>r</u>adiation)



### **Wavelength Selectors**

Filters

Monochromators Grating – usually used Prism











# **Advantages of Grating Monochromators**

Wavelength independence of dispersion.

Fixed dispersion makes it easy to scan an entire spectrum at constant bandwidth after initial adjustment of slitwidth.

Better dispersion for same size of dispersing element.

Can disperse radiation in far UV and infrared regions where absorption prevents use of prisms.

### **Disadvantages of Grating Monochromators**

Produce great amounts of stray radiation.

Produce more high-order spectra.

>Both of these disadvantages can be minimized with filters.



### **Types of Photon Detectors**

<u>Photovoltaic Cells</u> (or <u>Barrier-Layer Cells</u>) – the radiant energy generates a current at the interface of a semiconductor layer and a metal.

<u>Phototubes</u> – radiation causes emission of electrons from a photosensitive solid surface.

<u>Photomultiplier Tubes</u> – contain a photoemissive surface as well as several additional surfaces that emit a cascade of electrons when struck by electrons from the photosensitive area.





### **Types of Photon Detectors (continued)**

<u>Photoconductivity Detectors</u> – absorption of radiation by a semiconductor produces electrons and holes, thus leading to enhanced conductivity.

<u>Silicon Photodiodes</u> – Protons increase the conductance across a reversed-biased pn junction. Used as <u>diode array detectors</u> to observe the entire spectrum simultaneously.







#### **Advantages of Fourier Transform Spectroscopy**

<u>Fellgett Advantage</u> – all of the resolution elements for a spectrum are measured simultaneously, thus reducing the time required to derive a spectrum at any given signal-to-noise ratio.

<u>Jacquinot Advantage</u> – the large energy throughput of interferometric instruments (which have few optical elements and no slits to attenuate radiation.

High wavelength precision, making signal averaging feasible.

Ease and convenience that data can be computer-manipulated.